

- approximation using  $R_k \approx R(kh)$ , with time-step h = T/N
- assume constant velocity along each discrete segment

 $\dot{R}(t) \approx (R_{k+1} - R_k)/h, \quad t \in [kh, (k+1)h].$ 

▶ Lagrangian approximated (for  $t \in [kh, (k+1)h]$ ) according to

$$L(R(t), \dot{R}(t)) \approx \frac{1}{2} \operatorname{tr} \left( \Lambda \left( \frac{R_{k+1} - R_k}{h} \right)^T \left( \frac{R_{k+1} - R_k}{h} \right) \right)$$

Using properties of the trace this simplifies to

$$L(R(t),\dot{R}(t))\approx -\frac{1}{h^2}\operatorname{tr}\left(\Lambda R_k^T R_{k+1}\right), \quad t\in [kh,(k+1)h].$$

#### Integrator derivation

Taking varitions with respect to each  $R_k$ 

$$\operatorname{tr}(\Lambda R_{k-1}^{\mathsf{T}} \delta R_k + \Lambda \delta R_k^{\mathsf{T}} R_{k+1} - \lambda_k R_k^{\mathsf{T}} \delta R_k) = 0,$$

or equivalently

 $\operatorname{tr}((\Lambda W_{k-1} + \Lambda W_k^T - \lambda_k)\eta_k) = 0,$ (1)where  $W_k = R_k^T R_{k+1}$  and  $\eta_k = R_k^T \delta R_k$ .

$$\lambda_k = \Lambda W_{k-1} + \Lambda W_k^T.$$

Since  $\lambda_k$  is symmetric

 $\Lambda W_{k-1} + \Lambda W_k^T = W_{k-1}^T \Lambda + W_k \Lambda,$ and if we define  $\mu_k = W_k \Lambda - \Lambda W_k^T$ , this becomes equivalent to  $\mu_k = W_{k-1}^T \mu_{k-1} W_{k-1}.$ Since  $\mu_k$  is skew-symmetric  $\Rightarrow \mu_k \in \mathfrak{so}(3)^*$ , this can be written as

$$\mu_k = \operatorname{Ad}_{W_{k-1}}^* \mu_{k-1}.$$

- (Marsden, Pekarsky, Shkoller)
- The discrete trajectory must satisfy

$$\delta \sum_{k=0}^{N-1} -\frac{1}{h} \operatorname{tr}(\Lambda R_k^T R_{k+1}) + \sum_{k=0}^N \frac{1}{2h} \operatorname{tr}(\lambda_k (R_k^T R_k - \mathsf{I})) = 0,$$

where  $\lambda_k \in \mathbb{R}^{3 \times 3}$  is a matrix of multipliers enforcing orthogonality of R. Variations of  $R_k$  and  $\lambda_k$  are free.

Note that

$$\operatorname{tr}(\lambda_k(R_k^T R_k - \mathsf{I})) = \operatorname{tr}(\lambda_k^T(R_k^T R_k - \mathsf{I}))$$

then a symmetric matrix  $\lambda_k$  is sufficient to enforce this condition.

### General integrator equations

In summary, the equations of the integrator are

$\mu_k = \operatorname{Ad}_{W_{k-1}}^* \mu_{k-1},$	% Momentum update: explicit	(2)
$\mu_k = W_k \Lambda - \Lambda W_k^T,$	% Legendre transform: implict	(3)
$W_k^T W_k = I,$	% Orthogonality constraint: implict	(4)
$R_{k+1}=R_kW_k.$	% Rotation update: explicit	(5)

Given:  $\mu_{k-1}, W_{k-1}, R_k$ 

First update  $\mu_k$  from  $\mu_{k-1}$  and  $W_{k-1}$  (explicitly); then find  $W_k$ from  $\mu_k$  (implicitly); update  $R_{k+1}$  from  $R_k$  and  $W_k$  (explicitly).



Can we accurately reproduce these periodic

The sphere with radius  $\|\mu_0\|$ 

## **Direct Solution**

Solve for W in terms of  $\mu$  directly using the matrix elements  $W^{ij}$  as uknowns. The equation  $\mu = W\Lambda - \Lambda W^T$  is equivalent to requiring that

$$\mu_{1} = \Lambda_{2} W^{32} - \Lambda_{3} W^{23},$$
  

$$\mu_{2} = \Lambda_{3} W^{13} - \Lambda_{1} W^{31},$$
  

$$\mu_{3} = \Lambda_{1} W^{21} - \Lambda_{2} W^{12}.$$
  
(6)

where

$$\mu = \left[ \begin{array}{ccc} 0 & -\mu_3 & \mu_2 \\ \mu_3 & 0 & -\mu_1 \\ -\mu_2 & \mu_1 & 0 \end{array} \right],$$

and  $\Lambda_i$  are the diagonal elements of the matrix  $\Lambda$ .

> 3 linear equations + 6 orthogonality conditions on W:
 9 implicit equations ⇒ polynomial roots / Newton method.

### Explicit approximate solution

Use  $\tau$  to represent W but  $\underline{\mathrm{trunacte}}$  its expression in the Legendre transform equation

First order truncation of the exponential map  $\exp(\xi) = \sum_{i=0}^{\infty} \frac{\xi^i}{i!}$ :

$$\tau(\widehat{\omega}) = \exp(\widehat{\omega}) \approx \mathsf{I} + \widehat{\omega}.$$

The Leg. equation reduces to μ = ŵΛ + Λŵ, and using standard inertia matrix J = diag(J<sub>1</sub>, J<sub>2</sub>, J<sub>3</sub>) in identification so(3) ~ ℝ<sup>3</sup>, the integrator is <u>explicit</u>:

$$\omega_k = J^{-1} W_{k-1}^T J \omega_{k-1}$$
$$W_k = \exp(\omega_k),$$
$$R_{k+1} = R_K W_k.$$

#### Continuous Equations of Motion

- ► Constraints  $\mathcal{D}_q$  defined by *m* functions  $\omega^a : TQ \to \mathbb{R}$ , a = 1, ..., m linear in the velocities and satisfy  $\omega^a(q, \dot{q}) = 0$ .
- After taking variations we get

$$\dot{\boldsymbol{q}} = \boldsymbol{v},$$

$$\boldsymbol{p} = \frac{\partial \boldsymbol{L}}{\partial \boldsymbol{v}},$$

$$\langle \dot{\boldsymbol{p}} - \frac{\partial \boldsymbol{L}}{\partial \boldsymbol{q}} - \boldsymbol{f}, \delta \boldsymbol{q} \rangle =$$

$$\boldsymbol{\omega}^{a}(\boldsymbol{q}) \cdot \boldsymbol{v} = \boldsymbol{0}.$$

0.

Allowed variations are such that  $\omega^a(q) \cdot \delta q = 0$  and

$$\dot{p} = \frac{\partial L}{\partial q} + f + f_{con},$$

where f<sub>con</sub> are forces necessary to enforce constraints.
 f<sub>con</sub> = λ<sub>∂</sub>ω<sup>∂</sup>(q) to cancel any acceleration in ṗ - ∂L/∂q - f not aligned with the constraints. λ<sub>∂</sub> are called Lagrangian multipliers denoting the magnitute of the constraint forces.

#### Parametrized Solution

Parametrize  $W = \tau(\xi)$  using parameters  $\xi$ : e.g. exponential coordinates, or Cayley parameters

• Lie algebra identification  $\mathfrak{so}(3) \sim \mathbb{R}^3$ . Define  $\widehat{\cdot} : \mathbb{R}^3 \to \mathfrak{so}(3)$ 

$$\hat{v} = \begin{bmatrix} 0 & -w_3 & w_2 \\ w_3 & 0 & -w_1 \\ -w_2 & w_1 & 0 \end{bmatrix}$$
(7)

 $\mathfrak{so}(3)$  basis  $\{\widehat{e}_1, \widehat{e}_2, \widehat{e}_3\}, \widehat{e}_i \in \mathfrak{so}(3)$  where  $\{e_1, e_2, e_3\}$  is the standard  $\mathbb{R}^3$ -basis. Elements  $\xi \in \mathfrak{so}(3)$  correspond to  $\omega \in \mathbb{R}^3$  by  $\xi = \omega^{\alpha} \widehat{e}_{\alpha}$ , or  $\xi = \widehat{\omega}$ . Operator Ad becomes  $\operatorname{Ad}_R \omega = R\omega$ .

Example: use the <u>Cayley</u> map

$$\operatorname{cay}(\widehat{\omega}) = 1 + \frac{4}{4 + \|\omega\|^2} \left(\widehat{\omega} + \frac{\widehat{\omega}^2}{2}\right).$$
(8)

 $\blacktriangleright$  In order to compute W given  $\mu$  we solve

$$\mu = \mathsf{cay}(\widehat{\omega}) \mathsf{\Lambda} - \mathsf{\Lambda}\,\mathsf{cay}(-\widehat{\omega})$$

for the 3 elements of  $\omega$  (implicitly) and then find  $W = \tau(\hat{\omega})$ .

## Nonholonomic Dynamics

Systems with nonintegrable constraint on the velocities

- ► configuration space Q
- ▶ regular distribution  $\mathcal{D}$ : collection of subspaces  $\mathcal{D}_q \subset \mathcal{T}_q Q$
- Lagrangian  $L: TQ \rightarrow \mathbb{R}$ ,
- control force  $f : [0, T] \rightarrow T^*Q$ .

For a curve (q(t), v(t), p(t)) in  $TQ \oplus T^*Q$ ,  $t \in [0, T]$  the d'Alembert-Pontryagin principle states that

$$\begin{split} & \int_{0}^{T} \left[ L(q, v) + \langle p, \dot{q} - v \rangle \right] dt + \int_{0}^{T} \langle f, \delta q \rangle dt = 0, \\ & \delta q \in \mathcal{D}_{q} \quad \text{and} \quad v_{q} \in \mathcal{D}_{q}, \end{split}$$

for variations that vanish at the endpoints. Note: nonholonomic vs. vakonomic  $(\delta q \in \mathcal{D}_q)$ symplectic structure not preserved in general

#### The vertical rolling disk

Configuration space  $Q = SE(2) \times S^1$ , with pose  $(x, y, \phi) \in SE(2)$  and the rotation angle  $\theta \in S^1$ .

Lagrangian

$$L(x, y, \phi, \theta, \dot{x}, \dot{y}, \dot{\phi}, \dot{\theta}) = \frac{1}{2}m(\dot{x}^2 + \dot{y}^2) + \frac{1}{2}I\dot{\theta}^2 + \frac{1}{2}J\dot{\phi}^2,$$

m: mass, I, J: moments of intertia

▶ nonholonomic constraints are (with *R*: the disk radius)

$$\dot{x} = R(\cos \phi) \dot{ heta}, \qquad \dot{y} = R(\sin \phi) \dot{ heta}$$

or in the form  $\omega^a \cdot (\dot{x}, \dot{y}, \dot{\phi}, \dot{\theta}) = 0$ 

$$\omega^1 = (1, 0, 0, -R \cos \phi), \qquad \omega^2 = (0, 1, 0, -R \sin \phi)$$

• Controlled by torques  $u^{\phi}$  and  $u^{\theta}$ 

The vertical rolling disk: continuous equations

The momentum is

$$p = rac{\partial L}{\partial v} = (m\dot{x}, m\dot{y}, I\dot{\phi}, J\dot{ heta}).$$

The dynamics equation becomes

$$\begin{split} m\ddot{x} &= \lambda_{1}, \\ m\ddot{y} &= \lambda_{2}, \\ I\ddot{\phi} &= u^{\phi}, \\ J\ddot{\theta} &= -R\cos\phi\lambda_{1} - R\sin\phi\lambda_{2} + u^{\theta}. \end{split}$$

Differentiating the constraints and substituting

$$J\phi = u^{\phi},$$
$$(I + mR^2)\ddot{\theta} = u^{\theta},$$

which along with the constraints determine the dynamics.

# Discrete Equation of Motion

The resulting equations are

$$\begin{split} q_{k+1} &= q_k + hv_k, & \% \text{ configuration update} \\ p_k &= \frac{\partial L}{\partial v} (q_{k+\alpha}, v_k), & \% \text{ Legendre transform} \\ \frac{p_k - p_{k-1}}{h} & \% \text{ momentum update} \\ &= (1 - \alpha) \left( \frac{\partial L}{\partial q} (q_{k+\alpha}, v_k) + f_{k+\alpha} \right) + \alpha \left( \frac{\partial L}{\partial q} (q_{k-1+\alpha}, v_{k-1}) + f_{k-1+\alpha} \right) \\ &+ (\lambda_a)_k \omega^a (q_k), \\ \omega^a (q_{k+\alpha}) \cdot v_k &= 0. & \% \text{ velocity constraint} \end{split}$$

#### There is a lot more...

We've looked at only the most basic nonholonomic case. Things get more interesting when

- incorporating group symmetries (the principal bundle case)
- systmes with multiple bodies or internal joints
- studying relation to the continuous case (e.g. what is discrete connection, discrete curvature, discrete gyroscopic forces, how is momentum evolution different?)



## Nonholonomic Discrete Mechanics

▶ Discretization: represent 
$$q : [0, T] \rightarrow Q$$
 by set  $\{q_0, ..., q_N\}$ 

- Approximation:  $q(kh) \approx q_k$ , where h = T/N is the time-step.
- Discrete d'Alembert-Pontryagin principle

$$\delta \sum_{k=0}^{N-1} [hL(q_{k+\alpha}, v_k) + \langle p_k, (q_{k+1} - q_k) - hv_k \rangle] + \sum_{k=0}^{N-1} h \langle f_{k+\alpha}, \delta q_{k+\alpha} \rangle = 0,$$
  
$$\delta q_k \in \mathcal{D}_{q_k} \quad \text{and} \quad v_k \in \mathcal{D}_{q_{k+\alpha}},$$

where  $\alpha \in [0, 1]$ : determines the interpolate quadrature point; notation:  $x_{k+\alpha} := (1 - \alpha)x_k + \alpha x_{k+1}$ .

Vertical Disk Integrator Disk velocity  $v = (v^x, v^y, v^{\phi}, v^{\theta})$ . The discrete momentum is

$$p_{k} = \frac{\partial L}{\partial v}(q_{k+\alpha}, v_{k}) = (mv_{k}^{x}, mv_{k}^{y}, lv_{k}^{\phi}, Jv_{k}^{\theta})$$

The discrete constraints are

$$v_k^{\mathsf{x}} = R\cos(\phi_{k+\alpha})v_k^{\theta}, \qquad v_k^{\mathsf{y}} = R\sin(\phi_{k+\alpha})v_k^{\theta}.$$

From the dynamics equation the multipliers can be computed as

$$\lambda_1 = m(v_k^x - v_{k-1}^x)/h, \qquad \lambda_1 = m(v_k^y - v_{k-1}^y)/h.$$

Substituting and simplifying the discrete dynamics becomes

$$\begin{split} &I(\mathbf{v}_{k}^{\phi} - \mathbf{v}_{k-1}^{\phi})/h = (1 - \alpha)u_{k-1+\alpha}^{\phi} + \alpha u_{k+\alpha}^{\phi} \\ &\left[ \left( J + mR^{2}\cos(\alpha h \mathbf{v}_{k}^{\phi}) \right) \mathbf{v}_{k}^{\theta} - \left( J + mR^{2}\cos((1 - \alpha)h\mathbf{v}_{k}^{\phi}) \right) \mathbf{v}_{k-1}^{\theta} \right]/h \\ &= (1 - \alpha)u_{k-1+\alpha}^{\theta} + \alpha u_{k+\alpha}^{\theta}. \end{split}$$